

## Questions

1 (IB)

- a) The star Wolf 359 has a parallax angle of 0.419 arcsecond.
- Describe how this parallax angle is measured.
  - Calculate the distance in light-year from Earth to Wolf 359.
  - State why the method of parallax can only be used for stars at a distance less than a few hundred parsec from Earth.

b) The ratio

$$\frac{\text{apparent brightness of Wolf 359}}{\text{apparent brightness of the Sun}} \text{ is } 3.7 \times 10^{-15}.$$

Show that the ratio

$$\frac{\text{luminosity of Wolf 359}}{\text{luminosity of the Sun}} \text{ is } 8.9 \times 10^{-4}.$$

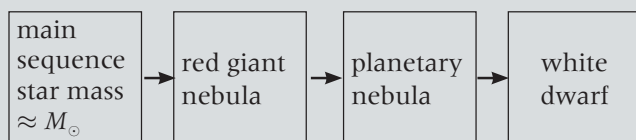
(11 marks)

- 2 The average intensity of the Sun's radiation at the surface of the Earth is  $1.37 \times 10^3 \text{ Wm}^{-2}$ . Calculate (a) the luminosity and (b) the surface temperature of the Sun.

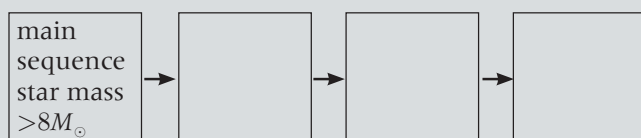
The mean separation of the Earth and the Sun =  $1.50 \times 10^{11} \text{ m}$ , radius of the Sun =  $6.96 \times 10^8 \text{ m}$ , Stefan–Boltzmann constant =  $5.67 \times 10^{-8} \text{ Wm}^{-2} \text{ K}^{-4}$ . (4 marks)

3 (IB)

The diagram below is a flow chart that shows the stages of evolution of a main sequence star such as the Sun. (Mass of the Sun, the solar mass =  $M_{\odot}$ )



- a) Copy and complete the boxes below to show the stages of evolution of a main sequence star that has a mass greater than  $8M_{\odot}$ .



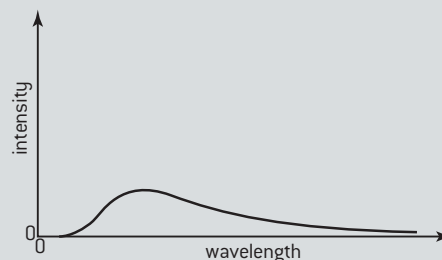
b) Outline why:

- white dwarf stars cannot have a greater mass than  $1.4M_{\odot}$ .
- it is possible for a main sequence star with a mass equal to  $8M_{\odot}$  to evolve into a white dwarf. (6 marks)

4 (IB)

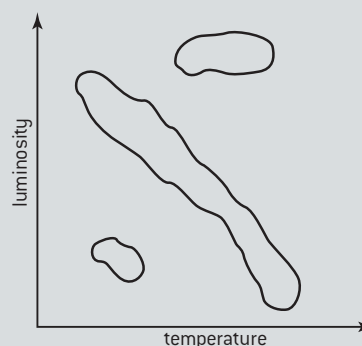
a) Define *luminosity*.

b) The sketch-graph below shows the intensity spectrum for a black body at a temperature of 6000 K.



On a copy of the axes, draw a sketch-graph showing the intensity spectrum for a black body at 8000 K.

c) A sketch of a Hertzsprung–Russell diagram is shown below.



Copy the diagram above and identify the:

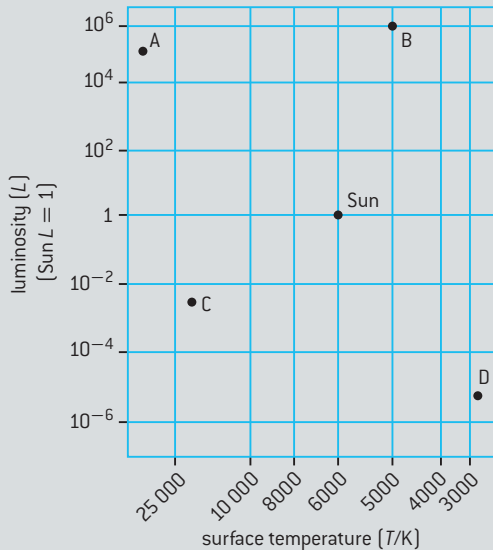
- main sequence (label this M)
- red giant region (label this R)
- white dwarf region (label this W).

d) In a Hertzsprung–Russell diagram, luminosity is plotted against temperature. Explain why the diagram alone does not enable the luminosity of a particular star to be determined from its temperature. (8 marks)



5 (IB)

The diagram below shows the grid of a Hertzsprung–Russell (HR) diagram on which the positions of the Sun and four other stars A, B, C and D are shown.



- Name the type of stars shown by A, B, C, and D.
- Explain, using information from the HR diagram and without making any calculations, how astronomers can deduce that star B is larger than star A.
- Using the following data and information from the HR diagram, show that star B is at a distance of about 700 pc from Earth.

Apparent brightness of the Sun =  $1.4 \times 10^3 \text{ W m}^{-2}$

Apparent brightness of star B =  $7.0 \times 10^{-8} \text{ W m}^{-2}$

Mean distance of the Sun from Earth = 1.0 AU

1 parsec =  $2.1 \times 10^5 \text{ AU}$  (11 marks)

6 (IB)

- State what is meant by *cosmic microwave background radiation*.
- Describe how the cosmic microwave background radiation provides evidence for the expanding universe. (5 marks)

7 (IB)

- In an observation of a distant galaxy, spectral lines are recorded. Spectral lines at these wavelengths cannot be produced in the laboratory. Explain this phenomenon.
- Describe how Hubble's law is used to determine the distance from the Earth to distant galaxies.
- Explain why Hubble's law is not used to measure distances to nearby stars or nearby galaxies (such as Andromeda). (6 marks)

8 (IB)

One of the most intense radio sources is the Galaxy NGC5128. Long exposure photographs show it to be a giant elliptical galaxy crossed by a band of dark dust. It lies about  $1.5 \times 10^7$  light years away from Earth.

- Describe any differences between this galaxy and the Milky Way.  
Hubble's law predicts that NGC5128 is moving away from Earth.
- (i) State Hubble's law.  
(ii) State and explain what experimental measurements need to be taken in order to determine the Hubble constant.
- A possible value for the Hubble constant is  $68 \text{ km s}^{-1} \text{ Mpc}^{-1}$ . Use this value to estimate:
  - the recession speed of NGC5128
  - the age of the universe. (10 marks)

9 a) Describe what is meant by a *nebula*.

- Explain how the Jeans criterion applies to star formation. (3 marks)

10 Outline how hydrogen is fused into helium in:

- stars of mass similar to that of the Sun
- stars of mass greater than ten solar masses. (6 marks)